

## HCO<sup>+</sup> and ionization structure observed by ALMA

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We observed HCO<sup>+</sup>(J=1–0) and H<sup>13</sup>CO<sup>+</sup> (J=1–0) emission towards the five protoplanetary disks around IM Lup, GM Aur, AS 209, HD 163296, and MWC 480 as part of the MAPS project[1][2]. HCO<sup>+</sup> is detected and mapped at 0.3'' resolution in all five disks, while H<sup>13</sup>CO<sup>+</sup> is detected (SNR >6 $\sigma$ ) towards GM Aur and HD 163296 and tentatively detected (SNR >3 $\sigma$ ) towards the other disks by a matched filter analysis. Inside a radius of  $R \sim 100$  au, the HCO<sup>+</sup> column density is flat or shows a central dip. At outer radii ( $\geq 100$  au), the HCO<sup>+</sup> column density decreases outwards, while the column density ratio of HCO<sup>+</sup>/CO is mostly in the range of  $\sim 10^{-5}$ – $10^{-4}$ . We derived the HCO<sup>+</sup> abundance in the warm CO-rich layer, where HCO<sup>+</sup> is expected to be the dominant molecular ion. At  $R \geq 100$  au, the HCO<sup>+</sup> abundance is  $\sim 3 \times 10^{-11}$ – $3 \times 10^{-10}$ , which is consistent with a template disk model with X-ray ionization. At the smaller radii, the abundance decreases inwards, which indicates that the ionization degree is lower in denser gas, especially inside the CO snow line, where the CO-rich layer is in the midplane. Comparison of template disk models with the column densities of HCO<sup>+</sup>, N<sub>2</sub>H<sup>+</sup>, and N<sub>2</sub>D<sup>+</sup> indicates that the midplane ionization rate is  $\geq 10^{-18}$  s<sup>-1</sup> for the disks around IM Lup, AS 209, and HD 163296. We also find hints of an increased HCO<sup>+</sup> abundance around the location of dust continuum gaps in AS 209, HD 163296, and MWC 480.

We also plan to show the summary of molecular column densities obtained in MAPS project and some model results for discussions in the poster session.

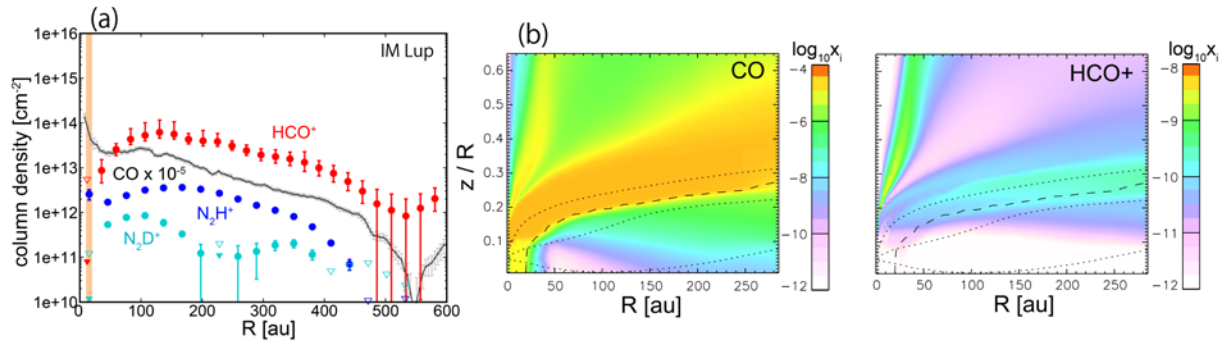


Figure: (a) Radial distributions of column densities of CO, HCO<sup>+</sup>, N<sub>2</sub>H<sup>+</sup>, and N<sub>2</sub>D<sup>+</sup> in IM Lup derived from MAPS project. (b) Distributions of CO and HCO<sup>+</sup> in a template disk model.

### References

- [1] Aikawa, Y., Cataldi, G., Yamato, Y. and MAPS team (2021) “Molecules with ALMA at Planet-forming Scales (MAPS) XIII: HCO<sup>+</sup> and disk ionization structure”, ApJS, in press
- [2] Oberg, K.I., Guzman, V.V., Walsh, C., Aikawa, Y., Bergin, E.A. and MAPS team (2021) “Molecules with ALMA at Planet-forming Scales (MAPS) I: Program Overview and Highlights”, ApJS in press